DG-SG method for Maxwell Equation

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1.DG method

For the simple example of partial differential equation,

$$\frac{\partial u}{\partial t} + \frac{\partial u}{\partial x} = f$$

First multiply both side by a test function v, and integrate it over the domain K,

$$\int_{K} \frac{\partial u}{\partial t} v + \int_{K} \frac{\partial u}{\partial x} v = \int_{K} f v$$

Then we can use the integration by parts formula,

$$\int_{K} \frac{\partial u}{\partial t} v + \int_{\partial K} \hat{u}v - \int_{K} u \frac{\partial v}{\partial x} = \int_{K} fv$$

Since we are using the orthonormal basis function ϕ , we can represent u, v, f by

$$u = \sum_i u_i \phi_i \quad v = \sum_i \phi_i \quad f = \sum_i f_i \phi_i$$

And we can denote $[u_1, u_2, ...]^T$ by u_h , and $[f_1, f_2, ...]^T$ by f_h

Since the basis function set is orthonormal, if $i \neq j$,

$$\int_K \phi_i \phi_j = 0$$

Then we can change

$$\int_{K} \frac{\partial u}{\partial t} v \quad to \quad \frac{d}{dt} I u_{h}, \qquad \int_{K} f v \quad to \quad I f_{h}$$

where I is the identity matrix

Here we use the central flux

$$\hat{u}=\frac{1}{2}(u^++u^-)$$

And we can also convert

$$\int_{\partial K} \frac{1}{2} (\phi_i^+ + \phi_i^-) \phi_j - \int_K \phi_i \frac{\partial \phi_j}{\partial x}$$

to matrix A using Gaussian quadrature formula

Then the partial differential equation can be changed to the linear multiplication,

$$\frac{d}{dt}u_h = -A*u_h + f_h$$

Hence we can use the explicit time advanced method

$$rac{u_h^{(t+1)} - u_h^{(t)}}{\Delta t} = -A * u_h^{(t)} + f_h$$

2. Sparse Grid

We consider the interval $\Omega = [0, 1]$, then the *n*-th level grid Ω_n consists of 2^n uniform cells

$$I_j^n = (2^{-n}j, 2^{-n}(j+1)), j = 0, ..., 2^n - 1$$

we can define

$$V_k^n := \{v : v \in P^k(I_j^n), \forall j = 0, ..., 2^n - 1\}$$

to be the usual piecewise polynomials of degree at most k. Then we have

$$V_k^0 \subset V_k^1 \subset V_k^2 \subset \dots$$

We can now define the multiwavelet subspace W_k^n , n = 1,2,...

$$V_k^{n-1} \bigoplus W_k^n = V_k^n, W_k^n \perp V_k^{n-1}$$

as the orthogonal complement of V_k^{n-1} in V_k^n with respect to the L^2 inner product

Now, for a multi-index $l = [l_1, l_2, ..., l_d] \in \mathbb{N}_0^d$, the l^1 and l^{∞} norms are defined as

$$|oldsymbol{lpha}|_1 := \sum_{m=1}^d lpha_m, |oldsymbol{lpha}|_\infty := \max_{1 \leq m \leq d} lpha_m$$

where d is the dimension and \mathbb{N}_0^d denotes the set of nonnegative integers

We define the tensor-product mesh grid

$$\Omega_l = \Omega_{l_1} \bigotimes \Omega l_2 \bigotimes ... \bigotimes \Omega_{l_d}$$

Based on the tensor product construction, we have

$$\mathbf{W}_{k}^{l} = W_{k,x_{1}}^{l_{1}} \times W_{k,x_{2}}^{l_{2}} \times ... \times W_{k,x_{d}}^{l_{d}}$$

Also based on the one-dimensional hierarchical decomposition, we have

$$m{V}_k^l = V_{k,x_1}^{l_1} imes V_{k,x_2}^{l_2} imes ... imes V_{k,x_d}^{l_d} = igoplus_{j_1 \leq l_1,...,j_d \leq l_d} m{W}_k^j$$

$$oldsymbol{V}_k^N = V_{k,x_1}^N imes V_{k,x_2}^N imes ... imes V_{k,x_d}^N = igoplus_{|l|_\infty \leq N} oldsymbol{W}_k^j$$

Where V_k^N is denoted if we use $h_N = 2^{-N}$ in each direction

And the sparse finite element approximation space \hat{V}_k^N is defined on Ω_N by

$$\hat{oldsymbol{V}}_k^N = igoplus_{|l|_1 \leq N} oldsymbol{W}_k^j$$

The degrees of freedom of sparse grid space scales as

$$\mathcal{O}((k+1)^d 2^N N^{d-1})$$

which is significantly less than that of traditional space with

$$\mathcal{O}((2^N(k+1))^d)$$

This is the key for computational savings and reduction in high dimensions.

3. Maxwell Equation

As for the Maxwell Equation,

$$rac{\partial m{B}}{\partial t} +
abla imes m{E} = 0$$
 $\epsilon \mu rac{\partial m{E}}{\partial t} -
abla imes m{B} = -\mu m{J}$

And for piecewise functions, we introduce the jumps and averages as follows

For any edge $e = K^+ \cap K^-$, with \mathbf{n}^{\pm} as the outward unit normal to ∂K^{\pm}

the jumps across e are defined as

$$[oldsymbol{U}]_{oldsymbol{n}} = oldsymbol{U}^+ \cdot oldsymbol{n}^+ + oldsymbol{U}^- \cdot oldsymbol{n}^- \ [oldsymbol{U}]_{ au} = oldsymbol{U}^+ imes oldsymbol{n}^+ + oldsymbol{U}^- imes oldsymbol{n}^-$$

and the averages are

$$\{m{U}\} = rac{1}{2}(m{U}^+ + m{U}^-)$$

Use the similar idea we introduced above

the semi-discrete DG methods for Maxwell equation are defined by following

$$\int_{K} \partial_{t} \boldsymbol{B}_{h} \cdot \boldsymbol{V} \ d\boldsymbol{x} = -\int_{K} \boldsymbol{E}_{h} \cdot \nabla \times \boldsymbol{V} \ d\boldsymbol{x} - \int_{\partial K} \widehat{\boldsymbol{n} \times \boldsymbol{E}}_{h} \cdot \boldsymbol{V} \ ds$$

$$\int_{K} \epsilon \mu \partial_{t} \boldsymbol{E}_{h} \cdot \boldsymbol{V} \ d\boldsymbol{x} = \int_{K} \boldsymbol{B}_{h} \cdot \nabla \times \boldsymbol{U} \ d\boldsymbol{x} + \int_{\partial K} \widehat{\boldsymbol{n} \times \boldsymbol{B}}_{h} \cdot \boldsymbol{U} \ d\boldsymbol{s} - \int_{K} \mu \boldsymbol{J} \cdot \boldsymbol{U} \ d\boldsymbol{x}$$

where $\boldsymbol{E}_h, \boldsymbol{B}_h, \boldsymbol{U}, \boldsymbol{V}$ are from the wavelet sparse grid basis

and all hat functions denote the numerical fluxes.

4. Numerical Flux

central flux:
$$\hat{E}_h = \frac{1}{2}(E_h^+ + E_h^-), \hat{B}_h = \frac{1}{2}(B_h^+ + B_h^-)$$

alternating flux:
$$\hat{\boldsymbol{E}}_h = {\boldsymbol{E}_h}^+, \ \hat{\boldsymbol{B}}_h = {\boldsymbol{B}_h}^- \ or \ \hat{\boldsymbol{E}}_h = {\boldsymbol{E}_h}^-, \ \hat{\boldsymbol{B}}_h = {\boldsymbol{B}_h}^+$$

$$upwind\ flux: \quad \hat{m{E}_h} = \{m{E}_h\} + rac{1}{2}[m{B}_h]_{ au},\ \hat{m{B}_h} = \{m{B}_h\} - rac{1}{2}[m{E}_h]_{ au}$$

Let $\boldsymbol{U} = \boldsymbol{E}_h$, $\boldsymbol{V} = \boldsymbol{B}_h$,

$$-\sum_{K} \int_{K} \mu \boldsymbol{J} \cdot \boldsymbol{E}_{h} \, d\boldsymbol{x} = \frac{1}{2} \frac{d}{dt} \sum_{K} \left(\int_{K} (\epsilon \mu |\boldsymbol{E}_{h}|^{2} + |\boldsymbol{B}_{h}|^{2}) \, d\boldsymbol{x} \right)$$
$$- \int_{\partial K} ([\boldsymbol{E}_{h} \times \boldsymbol{B}_{h}] + \hat{\boldsymbol{B}}_{h} \times [\boldsymbol{E}_{h}]_{\tau} - \hat{\boldsymbol{E}}_{h} \times [\boldsymbol{B}_{h}]_{\tau}) \, ds$$

And since for central flux and alternating flux,

$$[\boldsymbol{E}_h \times \boldsymbol{B}_h] + \hat{\boldsymbol{B}}_h \times [\boldsymbol{E}_h]_{ au} - \hat{\boldsymbol{E}}_h \times [\boldsymbol{B}_h]_{ au} = 0$$

Then we have energy conservation

$$-\sum_K \int_K \mu \boldsymbol{J} \cdot \boldsymbol{E}_h \ d\boldsymbol{x} = \frac{1}{2} \frac{d}{dt} \sum_K (\int_K (\epsilon \mu |\boldsymbol{E}_h|^2 + |\boldsymbol{B}_h|^2) \ d\boldsymbol{x}$$

And for upwind flux, the energy is decreasing,

$$-\sum_{K} \int_{K} \mu \boldsymbol{J} \cdot \boldsymbol{E}_{h} \, d\boldsymbol{x} = \frac{1}{2} \frac{d}{dt} \sum_{K} \left(\int_{K} (\epsilon \mu |\boldsymbol{E}_{h}|^{2} + |\boldsymbol{B}_{h}|^{2}) \, d\boldsymbol{x} + \frac{1}{2} \int_{\partial K} |[\boldsymbol{E}_{h}]_{\tau}|^{2} + |[\boldsymbol{B}_{h}]_{\tau}|^{2} \, ds \right)$$

Table 1: The accuracy of order for central flux

deg	level	$l-\infty$ error	order	$l-2 \ { m error}$	order
2	3	3.477E-03		9.847E-03	
2	4	1.999E-03	0.799	9.020E- 03	0.127
2	5	1.336E-03	0.581	7.767 E-03	0.216
2	6	5.509E-04	1.279	5.017 E-03	0.631
2	7	2.882 E-04	0.935	2.781E-03	0.851

Table 2: The accuracy of order for alternating flux

deg	level	$l-\infty$ error	order	$l-2 { m \ error}$	order
2	3	4.948E-03		2.195E-02	
2	4	3.293 E-03	0.587	1.529E-02	0.521
2	5	1.332 E-03	1.306	9.257 E-03	0.724
2	6	7.036E-04	0.921	5.265 E-03	0.814
2	7	2.626E-04	1.422	2.623E-03	1.006



